# Additional Results on the Ar V Spectrum

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Received April 12, 1994; revised manuscript received December 29, 1994

### I. Introduction

The ground state configuration of four times ionized argon,  $(Ar^{4+})$  is  $3s^23p^2$  with the terms  $^3P$ ,  $^1D$ , and  $^1S$ . The Ar V spectrum belongs to the Si I isoelectronic sequence. The spectra of the first and second elements in this sequence are presented in the book Atomic Energy Levels (AEL)<sup>[1]</sup>. A complete tabulation of the silicon energy levels including Si I levels was made by Martin and Zalubas<sup>[2]</sup>. Subsequent to this tabulation, Martin<sup>[3]</sup> and Svendenius et al.[4] reinvestigated and extended tlie knowledge of the spectrum of P II. Early results on tlie spectra of S III and Cl IV have been published in AEL. Lateroii, results for S III were published by Smitli et al. [5]. The Ar V spectrum was studied using various kinds of spectral sources. The spectra of argoii in tlie extreme ultraviolet were studied early by Boyce<sup>[6]</sup>, using an electrodeless light source, and by Phillips and Parker<sup>[7]</sup> using a spark source. Using a capillary source, Schönheit<sup>[8]</sup> studied the spectra of multiply ionized inert gases, including argon, and determined the degree

of ionization of the ions, but no classification of the lines was made. Some of these lines were later classified in tlie worlí of Ekberg et al.[9]. An experimental study of the low lying configurations in the Si I sequence was made by Smitt et al. [10]. Using a theta-pinch light source Fawcett et al.[11] have classified some lines of Ar V. In the worlí of De-Ye et al.[12] few lines of Ar V were also classified. Using the beam-foil technique Livingston et al.[13] have studied the argon spectra from Ar V to Ar VIII. Ellis and Martinson<sup>[14]</sup> predicted the  $3s3p^{3}$   ${}^{5}S_{2}^{o}$  level in the Si I sequence, that is the lowest energy level for quintet system. More recently, Trabert et al.[15-18] studied the VUV spectrum of argon and other ions in a search for iiitercombination lines of Mg, Al, and Si-lilíe ions. Recoil ions spectroscopy mas used by Lesteven-Vaïsse et al. [19] who have studied the argon spectra from Ar I to Ar X.

Many sets of theoretical calculations for the Si I isoelectronic sequence including Ar V levels, transition probabilities, oscillator strengths and lifetimes have heen published<sup>[22-25]</sup>. Experimental data for higher degrees of ionization in the Si I isoelectronic sequence were

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used by Biémont<sup>[20,21]</sup> (aiid references cited therein) to predict transitions from K VI through Ti IX, aiid energy levels and oscillator strengtlis from V X to Ni XV.

There is a renewed interest in spectroscopic data from rare gases due to applications in collision physics, laser physics, photoelectron spectroscopy and fusion diagnostics. In this last field the study of intercombination lines is important for the chiagnostics of laboratory and astrophysical plasmas<sup>[26,27]</sup>.

In the present work we report additional results for Ar V, applied to the study of the energy levels of the configuration  $3s^23p^2$ ,  $3s^23p3d$  and  $3s^24p4s$  and the transitions between them.

## III. Experiment

Tlie light source used in the present work is a discliarge tube built at Centro de Investigaciones Opticas (CIOp), La Plata, Argentina., to study highly ionized gases<sup>[28]</sup>. It is made with a pyrex tube with an initer dianieter of 3 mm and 300 nim long, with one end of the tube connected directly to a vacuum spectrograph through a higher aclaptor. The other end has a glass window for observations of discharge and alignment of the tube. The electrodes, placed 200 inm apart, were made of tungsten covered with indium. Gas excitation was produced by discharging through the tube a low inductance capacitor varied from 2.5 to 100 nF voltage up to 19 kV.

Light emitted axially was analyzed using a 3 ni normal incidence vacuum spectrograph with a concave diffraction grating with 1200 lines/mm, blazed for 1200 Å. The plate factor in the first order is 2.77 Å/mm. Ilford Q-2 plates were used to record the spectra. C, N, O, aiid ltnown lines of argon were also recorcled as internal wavelengthistandards. Exposing the plates with 10<sup>3</sup> shots we were able to obtain good lines of argon spectra. A number of experimental parameters, e. g., gas pressure, discharge voltage, and capacitance, were varied to distinguish among different states of ionization. A well developed Ar V spectrum was obtained with the following parameters: 175 mTorr, 18 kV, aiid 20 nF. Tlie positions of spectral lines on the plates were determined with a rotating prism photoelectric automatic

Grant comparator, with a precision of 1  $\mu$ m. The wavelength determined by this procedure is estimated to be  $\pm 0.01$  Å in the first diffraction order.

A similar esperiment using a tlieta-pinch built at tlie Instituto de Física at the Universidade Estadual de Campinas, Brazil<sup>[29]</sup>, confirmed tlie lines obtained in Argentina and was very lielpful in the task of assigning ionization stages to the individual spectral lines observed<sup>[31]</sup>. However we prefer to puhlish the lines of tlie first experiment because in Brazil tlie 2-m normal incidence spectrograph has a less favourable plate factor (4.61 Å/mm).

# III. Analysis

Line identifications were guided by theoretical predictions obtained from Cowan's computer codes<sup>[30]</sup>. The calculations were made at the Instituto de Física, UNICAMP. The preclictions were obtained by diagonalizing the energy matrices with appropriate relativistic Hartree-Fock (HFR) values for the energy parameters. The interpretation of configuration level structures was made by a least-square fit of the eliergy parameters to the observed levels.

Table I shows the 53 identified lines in the  $3s^23p^2 \leftarrow$  $3s3p^3$ ,  $3s^23p^2 \leftarrow 3s^23p3d$  and  $3s^23p^2 \leftarrow 3s^23p4s$  transition arrays. For 9 of these lines the classification is new. Table II shows tlie 24 determined energy levels belonging to the  $3s^23p^2$ ,  $3s3p^3$ ,  $3s^23p3d$  and  $3s^23p4s$ configurations with the uncertainty for each level. Optimization of thie energy level values was done from the observed wavelengtlis by an iterative procedure<sup>[32,33]</sup> in which the individual wavelengths are weighted according to their uncertainties, where 3 of tliese levels are new. Table III shows the energy parameters for tlie  $3ps^23p^2$  ground configuration. The standard deviation of the energy levels is 24 cm<sup>-1</sup> for 5 observed levels. Table IV shows the energy parameters for  $3s3p^3$ ,  $3s^23p3d$  aiid  $3s^23p4s$  configurations. The standard deviation of the fit is 151 cm<sup>-1</sup> for 23 observed levels. In order to obtain a better interpretation of the  $3s^23p^2$ ground configuration it was necessary to introduce in tlie calculation tlie  $3p^4$ ,  $3s3p^23d$  aiid  $3s^23d^2$  excited configurations. The  $3s^23p^2$  configuration is mainly affected

Table I. Classified lines in the  $3s^23p^2 \leftarrow 3s3p^3$ ,  $3s^23p^2 \leftarrow 3s^23p3d$  and  $3s^23p^2 \leftarrow 3s^23p4s$  transition arrays in Ar V.

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Intensity <sup>a</sup>	Wavelengtli (Å)		ransit	
1	336.57	$3s^23p^2 \stackrel{3}{_{2}} P_1$	-	$3s^2 3p4s \ ^3P_2^0$
1	337.58	$3s^23p^2 \ ^3P_0$	_	$3s^2 3p4s \ ^3P_1^0$
1	338.01	$3s^23p^2 \ ^3P_2$		$3s^2 3p4s \ ^3P_2^0$
1	338.45	$3s^2 3p^2 3P_1$	_	$3s^2 3p4s \ ^3P_1^0$
1	339.02	$3s^23p^2 \ ^3P_1$	_	$3s^2 3p4s \ ^3P_0^0$
1	339.89	$3s^2 3p^2 ^3 P_2$	-	$3s^2 3p4s \ ^3P_1^0$
2	350.88	$3s^23p^{2-1}D_2$	_	$3s^2 3p4s^{-1} P_1^0$
1	$357.23^{ m b}$	$3s^23p^{2-1}D_2$	-	$3s^2 3p4s \ ^3P_1^0$
3	379.69	$3s^23p^{2-1}S_0$		$3s^2 3p4s  ^1P_1^0$
1	$411.01^{\mathrm{b}}$	$3s^23p^{2-3}P_2$	-	$3s^23p3d^{-1}F_3^0$
3	436.63	$3s^23p^{2-1}D_2$	_	$3s^23p3d\ ^1F_3^0$
3	445.97	$3s^23p^2$ $^3P_0$		$3s^23p3d^{-3}D_1^0$
2	446.96	$3s^2 3p^2 \ ^3P_1$	_	$3s^23p3d\ ^3D_2^0$
3	447.53	$3s^23p^2$ $^3P_1$	_	$3s^23p3d\ ^3D_1^{\tilde{0}}$
2	449.08	$3s^23p^2 \ ^3P_2$	_	$3s^23p3d\ ^3D_3^{\hat{0}}$
3	449.50	$3s^23p^2$ $^3P_2$	_	$3s^23p3d\ ^3D_2^0$
3	452.39 <sup>b</sup>	$3s^2 3p^2 \ ^3P_1$	_	$3s^23p3d \ ^1D_2^0$
$\frac{\circ}{2}$	454.99 <sup>b</sup>	$3s^23p^2$ $^3P_2$		$3s^23p3d^{-1}D_2^0$
3	458.09	$3s^23p^2$ $^3P_0$	_	$3s^2 3p3d \ ^3P_1^0$
3	458.96	$3s^23p^2 \ ^3P_1$	_	$3s^2 3p3d^{-3} P_0^0$
3	459.73	$3s^{2}3p^{2}$ $^{3}P_{1}$	_	$3s^2 3p3d {}^3P_1^0$
$\frac{3}{2}$	461.24	$3s^{2}3p^{2} {}^{3}P_{1}$	_	$3s^2 3p3d {}^3P_2^0$
$\frac{z}{2}$	462.41	$3s^{2}3p^{2} {}^{3}P_{2}$	_	$3s 3p3a r_2$
		$3s^{2}3p^{2-3}P_{2}$		$3s^23p3d \ ^3P_1^0$
3	463.94	$3s^{-}3p^{-}P_2$	-	$3s^23p3d \ ^3P_2^0$
1	466.79	$3s^23p^2 {}^{1}S_0$		$3s^23p3d \ ^1P_1^0$
3	486.57 <sup>b</sup>	$3s^23p^2 \ ^1D_2$	-	$3s^23p3d^{-1}D_2^0$
1	495.09	$3s^23p^2 \ ^1D_2$	_	$3s^23p3d^{-3}P_1^0$
2	511.89	$3s^23p^2 \ ^3P_0$	_	$3s3p^{3-1}P_1^0$
2	513.90	$3s^23p^2 \ ^3P_1$		$3s3p^{3-1}P_1^0$
2	517.26	$3s^23p^2 \ ^3P_2$	_	$3s3p^{3} \stackrel{1}{}_{2}P_{1}^{0}$
3	522.08	$3s^23p^2 \ ^3P_0$	_	$3s3p^3 \ {}^3S_1^0$
3	524.19	$3s^23p^2 \ ^3P_1$	_	$3s3p^3 \ \ ^3S_1^0$
3	527.68	$3s^23p^2 \ ^3P_2$	_	$3s3p^3$ $^3S_1^0$
2	536.75	$3s^23p^{2-1}S_0$	_	$3s3p^3 \ ^3D_1^0$
3	558.48	$3s^23p^2 \stackrel{1}{\circ} D_2$		$3s3p^{3-1}P_1^0$
2	570.65	$3s^23p^2 \ ^1D_2$		$3s3p^3 \ ^3S_1^0$
4	635.17	$3s^23p^2 \ ^1S_0$		$3s3p^{3-1}P_1^0$
2	$650.95^{\mathrm{b}}$	$3s^23p^{2-1}S_0$	-	$3s3p^3 \ ^3S_1^0$
3	651.68	$3s^23p^2$ $^3P_1$	_	$3s3p^{3-1}D_2^0$
4	705.33	$3s^23p^2 \ ^3P_0$	_	$3s3p^{3-3}P_1^0$
2	$708.57^{ m b}$	$3s^23p^2 \ ^3P_1$		$3s3p^{3-3}P_0^0$
5	709.20	$3s^23p^2$ $^3P_1$	_	$3s3p^3 \ ^3P_2^0$
5	715.64	$3s^23p^2 ^3P_2$	_	$3s3p^{3-3}P_2^0$
4	725.09	$3s^23p^{2-1}D_2$	_	$3s3p^{3-1}D_2^0$
4	822.17	$3s^23p^2 ^3P_0$	-	$3s3p^{3-3}D_1^0$
4	827.03	$3s^23p^2$ $^3P_1$		$3s3p^{3-3}D_2^0$
4	827.35	$3s^23p^2$ $^3P_1$	-	$3s3p^{3-3}D_1^0$
4	834.91	$3s^23p^2$ $^3P_2$	_	$3s3p^3 \ ^3D_3^0$
$\overline{4}$	835.79	$3s^23p^2$ $^3P_2$		$3s3p^{3-3}D_2^0$
$\overline{4}$	836.13	$3s^23p^2$ $^3P_2$	_	$3s3p^3 \ ^3D_1^0$
$\overset{-}{2}$	$948.98^{b}$	$3s^23p^2 \ ^1D_2$	_	$3s3p^3 \ ^3D_2^0$
2	949.38 <sup>b</sup>	$3s^23p^2 \ ^1D_2$	_	$3s3p^{3-3}D_1^0$
	0.10,00	p D2		000p D1

<sup>&</sup>lt;sup>a</sup>The intensities of tlie lines are visual estimates of plate blackening. Tliey vary from 1 to 5. <sup>b</sup>New identification.

Table II. Energy levels of the  $3s^23p^2$ ,  $3s3p^3$ ,  $3s^23p3d$  and  $3s^23p4s$  configurations of Ar V

	Term Designation	Energy	Uncertainty
Configuration	C	$cm^{-1}$	$cm^{-1}$
	${}^{3}P_{0}$	0.0	2.0
	${}^{3}P_{1}$	767.1	2.0
$3s^23p^2$	${}^{3}P_{1}$	2032.1	1.0
	$^1D_2$	16302.0	1.0
	$^{1}S_{0}$	37919.4	2.0
	_		
	$^{3}D_{1}$	121631.8	1.0
	${}^{3}D_{2}$	121678.8	1.0
	$^{3}D_{3}$	121803.9	2.0
	${}^{3}P_{2}$	141768.6	1.0
$3s3p^{3}$	${}^{3}P_{1}$	141782.1	3.0
	${}^{3}P_{0}$	$141879.5^{[a]}$	3.0
	$^{1}D_{2}$	$154216.2^{[a]}$	1.0
	${}^{3}S_{1}$	191545.4	1.0
	$^{1}P_{1}$	195356.8	1.0
-	$^3P_2$	217575.6	3.0
	${}^{_{3}}P_{1}^{_{2}}$	218290.0	$\frac{3.0}{3.0}$
	${}^{1}P_{0}^{1}$	218651.1	14.0
	$^1D_2^0$	$221817.5^{[a]}$	3.0
$3s^23p3d$	$^3D_1$	224223.5	4.0
os opou	$^{3}\stackrel{D_{1}}{D_{2}}$	224501.3	8.0
	${}^{3}D_{3}^{2}$	224710.3	10.0
	$^{1}F_{3}$	$245335.9^{[a]}$	6.0
	${}^{1}P_{1}$	252135.0	5.0
	~ 1		0.0
	$^{3}P_{0}$	295735.0	9.0
$3s^23p4s$	${}^{3}P_{1}$	296234.9	5.0
-	$^3P_2^-$	297882.2	6.0
	${}^{1}P_{1}$	301295.7	6.0

<sup>[</sup>a] New level determined in this work.

Table III. Energy Parameters for the  $3s^23p^2$  configurations of Ar V

Configuration	Parameter	HFR Value (cm <sup>-1</sup> )	Fittecl $\overline{\text{Value}^{[a]}}$ $(\text{cm}^{-1})$	ratio Fitted/HF
$\frac{1}{3s^23p^2}$	$E_{av}$	0	$26254 \pm 13$	
-1	$F^{2}(3p,3p)$	75818	$62651 \pm 53$	0.83
	$\zeta_{3p}$	1246	$1378 \pm 22$	1.11

[a] The rms deviation of the fit is 24 cm<sup>-1</sup> for 5 observed levels.

Table IV. Energy Parameters for the  $3s^23p^3$ ,  $3s^23p3d$  and  $3s^23p4s$  configurations of Ar V.

configuration	parameter	HFR Value	Fitted Value <sup>[a]</sup>	ratio
Comingulation	parameter			
		$(cm^{-1})$	$(cm^{-1})$	Fitted/HF
$3s3p^{3}$	$E_{av}$	136214	$164465 \pm 80$	1.21
	$F^{2}(3p, 3p)$	75647	$65985 \pm 283$	0.87
	$G^{1}(3s, 3p)$	102393	$93546 \pm 144$	0.91
	$\zeta_{3p}$	1246	$1395 \pm 101$	1.12
$3s^23p3d$	$E_{av}$	198613	$221995 \pm 122$	1.12
	$F^2(3p, 3d)$	69452	$62913 \pm 364$	0.91
	$G^1(3p, 3d)$	86465	$74797 \pm 399$	0.86
	$G^{3}(3p, 3d)$	53372	$40386 \pm 415$	0.76
	$\zeta_{3p}$	1267	$1419 \pm 103$	1.12
	$\zeta_{3d}$	$49^{[b]}$		
$3s^23p4s$	$E_{av}$	289964	$298223 \pm 81$	1.13
	$G^{1}(3p, 4s)$	7978	$6516 \pm 109$	0.82
	$\zeta_{3p}$	1349	$1510 \pm 109$	1.12

[a] The rms deviation of the fit is 24 cm<sup>-1</sup> for 10 observed levels.

[b] Fixed during the fitting procedure.

by the  $3p^4$  configuration through the  $s^2 \leftarrow p^2$  interaction. The  $3s^23p^2$ ,  $3s3p^23d$  and  $3s^23d^2$  configurations are also interacting very strongly. In the calculations for the  $3s3p^3$   $3s^23p3d$  and  $3s^23p4s$  configurations we have included the  $3s^23p4d$ ,  $3p^33d$  and  $3s3p3d^2$  configurations to take into account their interaction.

## Acknowledgement s

We wish to thank Dr. Edison Shibuya for tlie use of tlie microdensitometer (FAPESP proc. 88-3885-0). This work lias been supported by Fundação de Amparo à Pesquisa do Estado de São Paulo (FAPESP), Brazil; Conselho Nacional de Desenvolvimento Científico e Tecnológico (CNPq), Brazil; Comission de Investigaciones Científicas de la Provincia de Buenos Aires (CIC), Argentina and Consejo Nacional de Investigaciones Científicas y Técnicas, Argentina.

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