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Elastic Scattering of 3 ~ and a Particles on Sb, Sn and Te\*

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The angular distributions of elastic scattering of 3 Haaend a particles of about 20 MeV from nuclei of Z approximately 50 were measured. Optical model parameters which reproduce the experimental data have been systematically determined.

Foram medidas as distribuições angulares do espalhamento elástico de <sup>3</sup>He rtículas a de cerca de 20 Mev por núcleos com Z da ordem de 5.0. Foram determinados sistematicamente os paramêtros do Modelo Optico que reproduzem os dados experirnentais.

### 1. INTRODUCTION

As part of a systematic study of <sup>3</sup>He and a particle induced reactions in the region of Tin, Antimony and Tellurium nuclei, the angular distributions of elastically scattered <sup>3</sup>He and a particles from these nuclei, were measured at the available incident energy of about 20 MeV. A compilation of <sup>3</sup>He and a particle optical model parameters fornuclei throughout the periodic table has been published by Perey and Perey<sup>1</sup>. More recent data and analysis for Z = 50 are available in Refs. (2) and (3). Nevertheless, at the incident energies available from the Pelletron accelerator, the data are scarce in the region of nuclei studied.

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# 2. EXPERIMENTAL TECHNIQUES AND RESULTS

# 2.1.- TARGETS AND APPARATUS

maare assuming that, at the incident energies utilized (-20 MeV), the cross section at tional Laboratory, Oak Ridge, Tennessee, USA The enrichment factors of the forward angles  $(\leq 45^{\circ})$  is only due to Rutherford scattering. The er-5%error evapora-Ridge Nasamples used to prepare the targets for the present experiments The target given in Table 1. The absolute cross sections were determined rors in the absolute cross sections are estimated to be 10%. A The targets of about 100 - 150 µg/cm² were prepared by vacuum teriais were obtained from the Stable Isotopes Division, Oak tion of enriched material onto 20  $\mu g/cm^2$  carbon foils. due to nonuniformity of the targets is included. the

°06 The beams of  ${}^{3}\mathrm{He}$  and a particles, produced in an external source of the (ME-20) into the Pelletron 8UD tandem accelerator of the Univerfocussing charge exchange type<sup>4</sup> , were injected through a single sity of São Paulo<sup>5</sup>. magnet

by After acceleration, the beam is energy analysed in a double focussing -eobtained measurements of (p,n) and ( $\alpha,n)$  thresholds ',' and the study od the The magnet calibration constant was . at 14231 keV in the  $^{12}C(p,p)$  reaction 90<sup>0</sup> magnet (ME-200). sonance

large the energy ത doublet, of analysed beam is transported to the entrance collimator aid of a switching magnet and a quadrupole scattering chamber. (m | = With the <u>e</u>

use of with solid targets, this chamber was installed and the centering tested After modifications for University The scattering chamber (Fig.1) was forrnely used at the Wisconsin to study the H(p,p)H reaction<sup>9</sup>. a precision better than 1%. ţ

105 A beam spot diameter of 4 mm could be obtained with the aid of the mm diameter 3.5 am long entrance collimator with defining slits of



Fig.1 - Vertical cass sectional view of the chamber. IS and IE defining slits of the beam entrance collimator; 2S and 2E antiscattering slits; 3. movesble arm; 4. target holder; 5. defining slits of the detector collimator; 6. antiscattering slits of the ortector collimator; 7. angle wheel; 8. connectors for detector signal output; 9. system for target sitioning; 10. Faraday cup; 11. electric supressor of Faraday cup; 12. detector.

The modified chamber allows for the changing of up to 4 targets without breaking the vacuum. Up to five silicon surface barrier detectors could be mounted on a movable arm and the positioning of the detectors could be read to a precision of 1'. The detector, provided with collimating and antiscattering slits, subtended a solid angle of  $-4 \times 10^{-4}$  steradian and had an angular resolution of  $\pm 1.4^{\circ}$ .

The beam current collected in a Faraday cup, which had geometrical and electrical suppression, was integrated to within a precision of 2%. A monitor detector, positioned at  $45^{\circ}$  to the beam, provided a constant check on the measurements.

Data were accumulated in a Honeywell DDP 516 Computer used in the multichannel mode and later transferred to a magnetic tape through an IBM / /360-44 Computer<sup>10</sup>.

# 2.2.- THE ELASTIC SCATTERING ANGULAR DISTRIBUTIONS

The angular distributions were measured from  $10^{\circ}$  to  $145^{\circ}$  in  $5^{\circ}$  steps for the following cases:

<sup>124</sup> Sn	( <sup>3</sup> He,	<sup>3</sup> He) <sup>124</sup> Sn,	22.35	MeV,
<sup>123</sup> Sb	(а,	$\alpha$ ) <sup>123</sup> Sb,	19.95	MeV,
<sup>124</sup> Te	(a,	a ) <sup>124</sup> Te,	19.3	MeV,
<sup>124</sup> Sn	(a,	a ) <sup>124</sup> Sn,	19.5	MeV,
<sup>123</sup> Sb	(³He,	<sup>3</sup> He) <sup>123</sup> Sb,	19.5	Me∨,
<sup>124</sup> Te	(³He,	<sup>3</sup> He) <sup>124</sup> Te,	19.52	MeV,
<sup>122</sup> Te	(³He,	<sup>3</sup> He) <sup>122</sup> Te,	19.52	MeV.

The statistical uncertainties varied from less than 1% in the forward angles to about 3% in the backward angles. In the case of <sup>124</sup>Sn + <sup>3</sup>He angular distribution, some of the backward angle data could not be obtained with statistical accuracy better than 14%.



Fig.2 - Typical spectrum of elastic scattering of  $^{12\,3}\text{Sb}(^{3}\text{He},^{3}\text{He})^{12\,3}\text{Sb}$  .

The energy resolution varied from 30 to 70 keV due to the energy loss of the beam in the target, kinematical broadening and detector resolution. An improvement of the energy resolution was obtained when a positive bias of 5 kV was applied to the target holder.

Particle identification was not considered necessary as the other reactions were of negligible cross section or were kinematically well separated from the elastic events.

Typical spectra for <sup>3</sup>He on <sup>123</sup>Sb and a on <sup>124</sup>Te are shown in Figs.2 and

On account of the rather low incident energy, the angular distributions do'not show any diffraction structure, but are rather smooth. The data reduced to the center of mass system are shown in Table 2.

# 3. OPTICAL MODEL ANALYSIS

# 3.1. - FORM OF THE OPTICAL POTENTIAL

The analyses of the experimental elastic angular distributioris were performed with the aid of an optical model search code  $MODOPT^{11}$  which could be run on both the IBM/360-44 and the B-6700 computers. The poreniial used was of the form:

$$\begin{split} U(r) &= V_{c}(r) - \{V_{0}f(r,r_{0},a_{0}) + i\left(W_{v}f(r,r_{v},a_{v}) + W_{s}g(r,r_{s},a_{s})\right)\} + \\ &+ \left(\frac{\hbar}{m_{\pi}c}\right)^{2}\{V_{so} \frac{1}{r}\frac{d}{dr}f(v,r_{so}^{R},a_{so}^{R}) + iW_{so}\frac{1}{r}\frac{d}{dr}f(r,r_{so}^{I},a_{so}^{I})\}^{\frac{1}{\sigma}} \cdot \hat{\ell} \end{split}$$

with

$$f(r, r_x, a_x) = \left[1 + \exp \frac{r - r_x A^{1/3}}{a_x}\right]^{-1}$$
 (Woods-Saxon),

$$g(r, r_x, a_x) = \exp\left(-\left\{\left(\frac{r-r_x A^{1/3}}{a_x}\right)\right\}^2\right)$$
 (Gaussian),

 $g(\mathbf{r}, \mathbf{r}_x, \mathbf{a}_x) = 4 a_x \frac{d}{d\mathbf{r}} f(\mathbf{r}, \mathbf{r}_x, \mathbf{a}_x)$  (Woods-Saxon derivative)

and

$$V_{c}(\mathbf{r}) = \frac{\frac{Z_{1} Z_{T} e^{2}}{2R_{c}} \left( 3 - \frac{r^{2}}{R_{c}^{2}} \right), \text{ for } \mathbf{r} < R_{c},$$
$$\frac{Z_{1} Z_{T} e^{2}}{\mathbf{r}} \text{ for } \mathbf{r} > R_{c},$$

where  ${\rm Z}_1$  and  ${\rm Z}_T$  are the atomic number of the incident particleandtarget respectively and

$$R_{c} = r_{c} A^{1/3} .$$

17 partial waves were necessary in the calculations. The Coulomb radius,  $\mathbf{r}_{c}$  was fixed at 1.25F as suggested by Luetzelschwab et  $al.^{12}$ .

### 3.2. - FOUR PARAMETER GRID SEARCH

In general, the optical potentials are determined with ambiguities. Both discrete and continuous ambiguities of the real potential are discussed by several authors  $^{13-16}$ .

There exist discrete values of the real well depth for a fixed radius which give almost the same fit (discrete ambiguity). The real potentials determined in the discrete ambiguity have the same depth at the radius of strongest absorption  $R_A$  which is related to the preferencial partial wave by the relation:

$$R_{A} = \frac{\eta + \{\eta^{2} + \ell (\ell+1)\}^{1/2}}{k}$$

where  $\eta$  is the Coulomb constant, k, the wave number and R is chosen so that  $\text{Re}(S_R) = 0.5$  (Ref.17), S being the scattering matrix.



Although the depths are completely different, for  $r > R_A$  all potentials have the same trend and similar fits to the elastic scattering data are obtained<sup>14,18,19</sup>. The discrete ambiguity corresponds to the inclusion of an additional half wave length of the wave function of the most significant partial wave into the potential well<sup>15</sup>.

The continuous ambiguity occurs for values of  $\mathbf{r}_0$  and  $V_0$  such that the product  $V_0 r_0^n$  is constant.

Two new parameters were introduced by Greenlees et  $al_{\bullet}^{20}$  to characterize the different potential families:  $J_0$ , thevolume integral per pair of particles,  $\langle r_0^2 \rangle^{1/2}$ , the root mean square radius.

For the Woods-Saxon form factor,

$$J = \frac{4\pi}{3} V_0 R_0^3 \left( 1 + \left(\frac{\pi \alpha_0}{R_0}\right)^2 \right)$$

with  $R_0 = r_0 A^{1/3}$  and  $J_0 = J/(A_1A_T)$ ,  $A_1$  and  $A_T$  are the mass numbers of the incident particle and of the target respectively.  $J_0$  varies in discrete steps of 100 MeV.F<sup>3</sup> for the different families<sup>19</sup>.

The root mean square radius for the Woods-Saxon form factor is given by,

$$< r_0^2 > \frac{1/2}{5} = \left(\frac{3}{5}R_0^2 \left\{1 + \frac{7}{5}(\pi \alpha_0/R_0)^2\right\}\right)^{1/2}$$

and should be constant as a function of V, within one family.

Based on a procedure suggested by Baugh<sup>15</sup>, a grid search on the parameters V, and **r**, was performed with the <sup>123</sup>Sb +  $\alpha$  data. Volume absorption was used in this procedure. No spin-orbit term was included. V, was varied from 40 to 240 MeV in steps of 20 Mev and  $r_0$  from 1.15 to 1.4 F in steps of 0.05 F. For each  $(V_0, r_0)$  selected pair, a simultaneous search was performed on the four parameters  $a_0$ ,  $W_2$ ,  $\mathbf{r}_V$  and  $a_2$ .



Fig.5- Variation of the optical model parameters as a function of  $V_0$  a) for  $^{12.3}$ Sb +  $^3$ He, b) for  $^{12.3}$ SB +  $\alpha$ .

The minimum  $\chi^2$  values were obtained with  $r_0 = 1.4$  F for values of V<sub>0</sub> at intervals of -80 MeV. For  $r_0 = 1.2$  F, even though the  $\chi^2$  minima were slightly worse, a spacing of -40 MeV between the minima was found. No visual distinction could be made between the fits obtained with  $r_0=1.2$  F and  $r_0 = 1.4$  F.

A four parameters grid search was also performed with the <sup>123</sup>Sb + <sup>3</sup>He data, for  $\mathbf{r}$ , = 1.2 F and 1.4 F. For  $\mathbf{r}$ , = 1.2 F, a V, spacing of -40 MeV between the  $\chi^2$  minima was obtained. However, for  $\mathbf{r}_0 = 1.4$  F, only one minimum with V, = 40 MeV was found.

These results are surnmarized in Fig.4, while the variation of the individual parameters as a function of the real depth V, for  $r_{r} = 1.2$  F. is presented in Fig.5. Starting from the parameters which gave minima for  $r_0 = 1.2$  F and 1.4 F in  ${}^{123}$ Sb( $\alpha, \alpha$ )  ${}^{123}$ Sb analysis, and for  $r_0 = 1.2$  F in <sup>123</sup>Sb(<sup>3</sup>He, <sup>3</sup>He) Sb analysis, a five parameter search was performed. The resulting parameters, together with the  $J_{n}$  and  $(r_{n}^{2})^{1/2}$  values are shown in Tables 3 and 4, for a particles and <sup>3</sup>He respectively. The parameters marked with (\*) in Table 3 and Table 4 were used as starting parameters for the other angular distributions. It is known that the best fits are obtained when the potential depth is 120-180 MeV for <sup>3</sup>He and 160-240 MeV for  $\alpha$  particles. With this criterion, the parameters given in Table 5 were chosen as the most adequate and the corresponding fits are shown in Fig.6.

### 3.3.- THE IMAGINARY PART OF THE OPTICAL POTENTIAL

Maintaining the parameters of the real potentials listed in Table 5, attempts were made to determine the surface absorption potentials (Woods-Saxon derivative form factor) for the Sn and Te angular distributions. In the case of Sb, a more detailed study was performed. The results are included in Table 3 and Table 4. No essential differences in the quality of the fits were observed between volume and surface absorption potentials, as expected, since the angular distributions are rather smooth<sup>14,18,19</sup>



**Fig.** 6 - Ratio of the experimental cross section to Rutherford cross section for elastic scattering of <sup>3</sup>He and a particles. The curves are calculated with the best fit parameters of Table 5. The error bars are shown when larger than the size of the point.

In the case of angular distributions with a marked diffraction pattern, a surface imaginary term gives, in general, a better fit than a volume term<sup>21</sup>.

### 3.4. - THE SPIN-ORBIT TERM

A spin orbit term of -2 MeV, as suggested by Luetzelschwab  $et \ all^{12}$ , Urone  $et \ all^{14}$  and by Cage  $et \ all^{19}$ , was included in the analysis without any significant improvement in the fit, as was expected for a structureless angular distribution. As no polarization data pertinent to the present results. were available, the spin-orbit term was not included in the final fit.

### 4. DISCUSSION AND CONCLUSIONS

An examination of Table 4 shows that the variation of the parameters follows the systematic behavior observed by Cage et  $al.^{19}$  and by Baugh<sup>15</sup>, namely: a, and  $r_v$  decrease whereas  $w_s$  or  $w_v$  and  $a_v$  increase for increasing  $V_o$ .

The radius of strongest absorption of the potentials (Table 4) which fit the <sup>3</sup>He scattering on Sb is 9.9 F and corresponds to R = 7, whereas for  $\alpha$  scattering on Sb (Table 3),  $R_A = 10.3$  F which corresponds to R = 9 (Fig.7). These values are in agreement with Cage *et al.*<sup>16</sup> and Weisser *et al.*<sup>18</sup> although the corresponding R values differ slightly from those calculated imposing the condition that Re( $S_p$ ) = 0.5 (Ref.17).

The values of  $V_0$  obtained in the grid search of  ${}^{123}Sb(\alpha, \alpha){}^{123}Sb$  are plotted as a function of r, in Fig.8. The mean value of  $J_0$  varies by about 100 MeV F for adjacent families

For some families, the valueç of  $\langle r_0^2 \rangle^{1/2}$  were calculated for different values of r. These results are shown in Fig.9. A slight variation of the root mean square radius, as a function of r, is observed in agreement with Cage et  $al^{1,19}$ , although it had been introduced as a constant within one family.

No preference for a surface or volume imaginary form factor was detected.

As a spin-orbit term was not necessary, effects due to the target spins were not looked for.

As can be seen in Fig. 6, all the *a* particle angular distributions (19.3 - 19.9 MeV) have the same overall trend. A similar behaviour is also observed for <sup>3</sup>He, except in the case of <sup>124</sup>Sn measured at a slightly higherenergy (22.3 MeV). The comparison between a particle and <sup>3</sup>He angular distributions measured at the same energy shows a larger diffuseness for a particle than for <sup>3</sup>He. Bock et  $al^{22}$  have observed a larger diffuseness for <sup>3</sup>He than for a particle with data obtained at equal momenta.

Systematic measurements at higher energies are needed to observe the dependence of optical model potential on more complicated effects such as mass number of targets, spin-spin interactions and isotopic spin.

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Target	Chemical Composition	%
<sup>122</sup> Te	metallic	96.21
<sup>124</sup> Te	metallic	96.21
<sup>124</sup> Sn	SnO <sub>2</sub>	94.74
<sup>123</sup> Sb	Sb0 <sub>2</sub>	99.05

Table 1 - Enrichment factors of the targets studied.



Fig.7 - Plots of the depth of real optical potentials of Tables 3 and 4 as a function of the real radius R .



Fig.8 - Variation of potential V as a function of  $r_0$  obtained in the search for  ${}^{123}Sb(\alpha,\alpha){}^{123}Sb$ . The corresponding  $\overline{J}_0$  values of each family are also shown.



Fig.9 - Variation of  $\langle x_0^2 \rangle^{1/2}$  as a function of **r**. The open circles correspond to the family with  $\overline{J}_0 = 260 \text{ MeV.F}^3$  and the crosses to  $\overline{J}_0 = 380 \text{ MeV.F}^3$ .

124Sn	( <sup>3</sup> He, <sup>3</sup> He) <sup>12</sup>	Sn E = 22.35	<sup>124</sup> Sn (a	,¢) <sup>124</sup> Sn	E= 19.5 MeV
<sup>0</sup> e1/a . 6 . X	e. x	10 <sup>2</sup>	e°	9-21-	eabs x 10 <sup>2</sup>
	Ca, Ruth.	abs.	Cm	el/oRuth.	205.
6	1.097	7.3	15.48	0.957	5.6
. 48	0.909	6.0	20.63	0.961	5.7
59	1.148	7.6	25.78	0.973	5.7
10	1.129	7.5	30.92	0.966	5.7
80	1.075	7.1	36.06	1.063	6.3
0	.866	5.7	41.19	1.072	6.2
9	.724	4.8	46.31	1.045	5.4
7	.619	4.2	51.42	.923	4.6
. 24	. 486	2.9	56.51	.786	3.8
21	.441	3.0	61.60	.650	3.2
26	. 344	2.4	66.68	.544	2.6
. 31	. 274	1.9	71.74	.441	2.1
. 34	.243	1.8	76.79	. 359	1.7
	205	1.4	91 92	286	1.4
	.203	1.4	01.02	234	1.1
	.142	1.0	00.04	. 234	1.1
• • •	.136	.93	91.85	.191	.94
. 3	9.10	.72	96.84	.154	./5
. 3	7.092	.69	101.82	.128	.65
5.3	5.0653	.62	106.79	.106	.55
1.3	1.064	. 53	111.74	.0905	.41
16.2	6.060	.60	116.68	.0750	.41
21.2	1.047	.42	121.60	.0668	.33
26.1	4 .042	.41	126.51	.0539	. 30
51.0	7.034	. 36	131.42	.0473	. 27
35.9	9 .0\$2	. 39	136.31	.0419	. 20
40.9	0.029	.27	141.19	.0316	.20
150.7	0.021	. 29	146.06	.0323	.21
160.4	8 .027	. 35			

Table 2 - Ratio of the measured elastic cross section to the Rutherford cross section and the corresponding absolute error.

12 Te (	<sup>3</sup> He, <sup>3</sup> He) <sup>12</sup> *Te	E = 19.52MeV	12*Te	(a.a) 12 Te	E = 19.3 MeV			
θ <sup>o</sup> cm	<sup>σ</sup> el/σ <sub>Ruth</sub> .	€ <sub>abs</sub> ,× 10 <sup>2</sup>	θcm	<sup>@</sup> el/σ <sub>Ruth</sub>	<sup>c</sup> abs. × 10 <sup>2</sup>			
15.36	0.961	4.6	15.48	1.037	5.0			6°cu
20.47	1.028	4.9	20.63	0.954	4.6			15.
25.59	1.018	4.8	25.78	1.055	5.1			20.
30.69	0.953	4.6	30.92	1.037	5.0			25.
35.80	0.991	4.7	36.06	0.955	4.6			30.
40.89	1.070	5.1	41.19	1.030	4.9			35.
45.98	1.049	5.0	46.31	1.047	5.1			40.
51.06	.880	4.2	51.42	1.000	4.8			46.
56.14	.815	3.9	56.51	.909	4.4			51.
61.20	.717	3.5	61.60	.814	3.9			56.
66.26	.640	3.1	66.68	.743	3.6			61.
71.30	.548	2.7	71.74	.639	3.1			65.
76.34	.476	2.3	76.79	.481	2.3			71.
81.37	.447	2.2	81.82	.405	1.9			76.
86.38	. 382	1.8	86.84	. 328	1.6			81.
91.39	.342	1.7	91.85	. 274	1.3			86.
96.38	.264	1.4	101.82	.195	.73	1		91.
101.37	.238	1.2	111.74	.151	.73			96.
106.34	.198	1.0	116.28	.117	.56			101.
111.30	.171	.87	121.60	.101	.50			106.
116.26	.148	.78	126.60	.0912	.51			111.
121.20	.146	.77	131.42	.0739	. 36			116.
126.14	.131	.72	136.31	.0624	. 37			121
131.06	.108	.55	146.06	.0528	. 33			126
135.98	.100	.50						131
140.89	.0872	.45						136
145.80	.0789	.45						140
								145

		<sup>122</sup> Te( <sup>3</sup> He, <sup>3</sup> He) <sup>122</sup> Te	E= 19.52 Me
	ecn	<sup>o</sup> el/c <sub>Ruth</sub> .	<sup>e</sup> abs,× 10 <sup>2</sup>
	15.36	0.961	4.1
	20.48	1.031	4.4
	25.60	1.027	4.4
	30.70	1.021	4.4
	35.81	1.003	4.3
	40.91	1.004	4.3
	46.00	.962	4.1
	51.08	.896	3.8
	\$6.15	.819	3.5
	61.22	.792	3.4
	65.28	.701	3.0
	71.32	.550	2.4
	76.36	.480	2.1
	81.39	.442	1.9
	86.40	.367	1.6
1	91.41	.305	1.3
	96.40	.265	1.2
	101.39	. 229	1.0
	106.36	.195	.88
	111.32	.172	. 77
-	116.28	.150	.67
	121.22	. 145	.65
İ	126.15	.129	.58
	131.08	.108	.54
	136.00	.0922	.46
1	140.91	.0849	.42
	145.81	.0819	.41

Table 2 - (Continued)

$r_{0} = 1.2 \text{ F}$ $r_{c} = 1.25 \text{ F}$												
Energy (MeV)	Element	V <sub>o</sub> (MeV)	a <sub>o</sub> (F	W <sub>v</sub> (MeV)	r <sub>v</sub> (]a <sub>v</sub>	(aF)	W <sub>s</sub> (MeV)	r <sub>s</sub> (F)	a <sub>s</sub> (F)	χ²	J <sub>o</sub> (MeV.F <sup>3</sup> )	$\left< r_{o}^{2} \right>^{1/2} (F)$
19.95	<sup>1 2 3</sup> Sb	50.44*	0.936	3.72	1.827	0.388				1.41	113	5.784
		99.87*	0.822	12.01	1.640	9538				1.28	214	5.540
		124.61*	0.733	7.46	1.708	Q 498		ļ		1.32	264	5.461
		176.29*	0.737	9.39	1.676	0.506		ļ		1.31	373	5.372
		237.90*	0.731	11.68	1.645	9201		1		1.30	489	5.305
		64.80	0.893				12.78	1.425	0.615	1.29		
		103.08	0.778				47.94	1.135	0.683	6.14		
		122.91	0.736			ł	17.64	1.407	0.591	1.30		
		173.85	0.745				24.09	1.397	0.555	1.27		
		249.78	0.646				15.93	1.252	0.787	3.20		
19.5	<sup>124</sup> Sn	42.07	0.952	3.87	1.831	0.532				0.22	95	5.816
		110.42	0.775	11.30	1.574	0.724				0.27	232	5.443
		124.42	0.796	9.29	1.630	0.517				0.63	264	5.484
		185.05	0.720	12.87	1.552	0.665				0.30	382	5.338
		239.22	0.684	12.45	1.534	0.719				0.57	487	5.272
		190.31	0.707				27.60	1.189	0.756	0.28		
19.3	<sup>12</sup> "Te	64.50	0.839	10.88	1.557	0.762	· ·			0.73	139	5.571
		125.92	0.729	9.88	1.552	0.786				0.72	261	5.355
		178.94	0.672	14.24	1.415	0.892		ĺ		0.71	363	5.251
		181.56	0.634				18.85	0.835	1.165	0.65		
	<u> </u>	<u> </u>			<b>.</b>							
r <sub>o</sub> =	.4 F	$r_c = 1.2$	5 F		,	,	·	-1		r	<b></b>	T
19.95	<sup>1 2 3</sup> Sb	77.9	0.658	7.62	1.701	0.464				1.≾0	243	5.920
		159.4	0.576	12.41	1.638	0.447				1.z4	488	5.801
		241.3	0.538	16.08	1.623	0.393				1.z4	733	5.750

Table 3 - Potential families obtained for  $\boldsymbol{\alpha}$  particles.

r = 1	.2 F	$r_c = 1.25$	$n_c = 1.25 \text{ F}$												
nergy (MeV)	Element	V <sub>o</sub> (MeV)	a <sub>o</sub> (F)	W <sub>v</sub> (MeV)	r <sub>v</sub> (F)	a <sub>ν</sub> (F)	W <sub>s</sub> (MeV)	r <sub>s</sub> (F)	a <sub>s</sub> (F)	x <sup>2</sup>	$J_{o}(MeV.F^{3})$	$\langle r_o^2 \rangle^{1/2} (F)$			
19.5	<sup>1 2 3</sup> Sb	83.48*	0.744	10.87	1.619	0.735				0.24	232	5.386			
	j	134.47*	0.686	13.01	1.561	0.760				0.24	367	5.276			
		196.24*	0.644	19.10	1.462	0.799	ļ			0.24	528	5.203			
	1	233.47*	0.617	22.15	1.421	0.816			1 1	0.24	623	5.157			
	1	84.51	0.747			l	17.21	1.260	0.809	0.25					
		134.11	0.688				25.11	1.182	0.816	0.25	1	1			
		194.85	0.630				26.02	1.108	0.881	0.25					
		232.94	0.626				39.57	1.065	0.842	0.25	ļ				
22.35	124 Sn	94.02	0.759	14.22	1.567	0.683				0.77	262	5.412			
		184.12	0.673	21.64	1.476	0.706				0.76	499	5.253			
		220.72	0.653	24.78	1.446	0.714				0.76	593	5.218			
		180.77	0.480				41.75	1.172	0.715	0.77					
19.5	<sup>124</sup> Te	66.45	0.734	7.53	1.716	0.762			1	0.67	183	5.364			
		142.74	0.551	7.50	1.502	1.154		ļ		0.64	372	5.054			
	1	205.15	0.557	12.48	1.340	1.174			1	0.64	536	5.063			
		172.02	0.624				20.03	1.070	1.002	0.79					
10 5	12270	87 92	0.754	8.08	1.684	0.735				0.70	245	5.414			
12.3	10	136.30	0.695	8.20	1.668	0.760			ł	0.69	372	5.304			
		108 77	0.660	11 74	1.568	0.700	ł			0.69	536	5.242			
		235 30	0.644	13.30	1.530	0.816	ļ			0.66	631	5.214			
		146 88	0 567	13.35	1	10.010	12.73	0.481	. 680	0.38					
	1	140.00	10.307					1.101							

Table 4 - Potential families obtained for  $^{3}$ He.

REACTION	E <sub>lab</sub> (MeV)	V <sub>o</sub> (MeV)	r <sub>o</sub> (F)	a <sub>o</sub> (F)	W <sub>v</sub> (MeV)	r <sub>v</sub> (F)	a <sub>v</sub> (F)	x <sup>2</sup>
<sup>124</sup> Sn( <sup>3</sup> He, <sup>3</sup> He) <sup>124</sup> Sn	22.35	184.12	1.2	0.673	21.64	1• 476	<b>0.</b> 706	0.76
$^{124}Sn(\alpha,\alpha)^{124}Sn$	19.5	185.05	1.2	0.720	12.87	1.552	0.666	0.30
<sup>123</sup> Sb( <sup>3</sup> He, <sup>3</sup> He) <sup>123</sup> Sb	19.5	134.47	1.2	0.686	13.01	1.561	0.760	0.24
$^{123}$ Sb( $\alpha$ , $\alpha$ ) $^{123}$ Sb	19.95	176.29	1.2	0.737	9.39	1.676	0.506	1.31
<sup>124</sup> Te( <sup>3</sup> He, <sup>3</sup> He) <sup>124</sup> Te	19.52	142.74	1.2	0.551	7.51	1.502	1.154	0.64
<sup>124</sup> Te( a , a ) <sup>124</sup> Te	19.3	178.94	1.2	0.672	14.24	1.415	0.892	0.71
<sup>122</sup> Te( <sup>3</sup> He, <sup>3</sup> He) <sup>122</sup> Te	19.52	136.39	1.2	0.695	8.29	1.668	0.760	0.69

Table 5 - Best fit parameters.

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